jtdtype.com © JTD, LLC 2021

#### Parabolica

Designed By James Hultquist-Todd 5 Weights 2 Optical Sizes Designed in 2023 Parabolica Hairline

Parabolica Hairline Oblique

Parabolica Regular

Parabolica Regular Oblique

Parabolica Medium

Parabolica Medium Oblique

**Parabolica Bold** 

Parabolica Bold Oblique

**Parabolica Black** 

Parabolica Black Oblique

Parabolica Text Regular

Parabolica Text Regular Oblique

Parabolica Text Medium

Parabolica Text Medium Oblique

**Parabolica Text Bold** 

Parabolica Text Bold Oblique

**Parabolica Text Black** 

Parabolica Text Black Oblique

#### k = 75 pci (20 MN/m3)DIGITAL AUDIO TAPE K-Factor Maximum 2% DIGITAL AUDIO TAPE **Negative Sync Pulses** INTERMODULATION **Entry Monitor System VOLTAGE SELECTOR Level Sensors Power** INJECTION-MOLDE

#### Single-Ended Triode IU/SLA ALIGNMENT 30 Hz-16 kHz ±3 dB PHOTOMULTIPLIER 30 Hz-16 kHz ±3 dB 10V PEAK-TO-PEAK **Crosswind Velocity** LEFT STATIC PORT

#### Luminance RECTIFIER

### Sharpless Epoxidation PCM/CCS ANTENNA

Pneumatic checkout racks regulates controls EXPONENTIAL FREQUENCY MODULATION

#### Declination RECTIFIER

### Thermal Conditioning MODULATED SIGNAL

See Volume 2 For Pin Voltages And Parts List EXPONENTIAL FREQUENCY MODULATION

### Luminance OPTACON

#### Ceramic Ablative Liner ELECTRODYNAMICS

Hydraulic supply and checkout unit gimbals 95 SECOND MAXIMUM START ENVELOPE

#### Magentizer THERMAL

#### Thermal Conditioning ELECTROCHEMICAL

Pneumatic checkout racks regulates controls ENSEMBLE DE LANCEMENT ARIANE-ELA4

## Calibration INVERTER

#### Wein Bridge Oscillator INTERMODULATION

Solid State FM Stereo Tuner Model TU-355 SYSTEM IS AS SHOWN FOR LINK P-1 ONLY

# Composite OPTACON

### Sharpless Epoxidation VOLTAGE SELECTOR

Propeller Model Number: 1C160/DTM7557 SYSTEM IS AS SHOWN FOR LINK P-1 ONLY

#### Resistance RECEIVER

#### Entry Monitor System CARBURATOR ICING

Stage Engine Actuator Measuring Voltages FOR .241 VOLT VTVM READING AT 500 \( \sqrt{} \)

## Thyratrons OPTACON

### 90° Oriented Ablative VOLTAGE SELECTOR

Hydraulic supply and checkout unit gimbals
95 SECOND MAXIMUM START ENVELOPE

### Attenuator VOCODER

### Launch Pad A, LC-39 16-BIT RESOLUTION

S/C Separation Shaped Charge Igniter Nº2
DIELECTRIC WITHSTAND VOLTAGE TEST

# Amplitude VOCDER

### **Unilateral Transducer INJECTION-MOLDED**

Attention, Risque De Choc, Ne Pas Enlever AT-F3 MOVING COIL PHONO CARTRIDGE

#### Waveform TETRODE

### Lossily Compression IU/SLA ALIGNMENT

Do not expose this unit to rain or moisture DO NOT BLOCK VENTILATION GRILLES

#### Alternator TETRODE

#### Single-Ended Triode RF MULTICOUPLER

Typical Inertial Azimuth During TLI Boost DO NOT BLOCK VENTILATION GRILLES

# Trajectory PHASING

### Không Được Đi Qua HATCH ACTUATOR

1/4 Phone or 3-Pin XLR Input Connectors MIKROELEKTRONIKAI VÁLLALAT / MEV

# Trajectory PHASING

#### Parasitic Oscillation RF MULTICOUPLER

Maximum Proportional Rate Command MULTIPLEXER ASSEMBLY MODEL 270

# Capacitor PHASING

#### Retrograde Section RF TRANSMISSION

Vent & Relief ValveThrust Vector Control MULTIPLEXER ASSEMBLY MODEL 270

# Ferrofluid DEGAUSS

#### Parasitic Oscillation LEFT STATIC PORT

Maximum Proportional Rate Command TAPE TENSION ADJUSTMENT POINTS

# Ferrofluid TETRODE

### Retrograde Section RF TRANSMISSION

Maximum Proportional Rate Command FEEDBACK CONTROL VOLTAGE INPUT

#### Capacitor TETRODE

#### Retrograde Section F-1 ENGINE VALVES

KSY34 High-Frequency NPN Transistor
TAPE TENSION ADJUSTMENT POINTS

There is an occasional star, like chi Carinae, whose spectrum consists almost wholly of bright lines, in general bearing no apparent relationship to the bright lines in the spectra of the gaseous nebulae except that the hydrogen lines are there, as they are almost everywhere. There is reason to believe that such a spectrum indicates the existence of a very extensive and very hot atmosphere surrounding the main body, or core, of the star in question. This particular star is remarkable in that it has undergone great changes in brilliancy and is located upon a background of nebulosity. The chances are strong that the star has rushed through the nebulosity with high rate of speed and that the resulting bombardment of the star

16 Pt

There is an occasional star, like chi Carinae, whose spectrum consists almost wholly of bright lines, in general bearing no apparent relationship to the bright lines in the spectra of the gaseous nebulae except that the hydrogen lines are there, as they are almost everywhere. There is reason to believe that such a spectrum indicates the existence of a very extensive and very hot atmosphere surrounding the main body, or core, of the star in question. This particular star is remarkable in that it has undergone great changes in brilliancy and is located upon a background of nebulosity. The chances are strong that the star has rushed through the nebulosity with high rate of speed and that the resulting bombardment of the star has expanded and intensely heated its atmosphere.

There are the Wolf-Rayet stars, named from the French astronomers who discovered the first three of this class, whose spectra show a great variety of combinations of continuous spectrum and bright bands. We believe that the continuous spectrum in such a star comes from the more condensed central part, or core, and that the

There is an occasional star, like chi Carinae, whose spectrum consists almost wholly of bright lines, in general bearing no apparent relationship to the bright lines in the spectra of the gaseous nebulae except that the hydrogen lines are there, as they are almost everywhere. There is reason to believe that such a spectrum indicates the existence of a very extensive and very hot atmosphere surrounding the main body, or core, of the star in question. This particular star is remarkable in that it has undergone great changes in brilliancy and is located upon a background of nebulosity. The chances are strong that the star has rushed through the nebulosity with high rate of speed and that the resulting bombardment of the star

16 Pt

There is an occasional star, like chi Carinae, whose spectrum consists almost wholly of bright lines, in general bearing no apparent relationship to the bright lines in the spectra of the gaseous nebulae except that the hydrogen lines are there, as they are almost everywhere. There is reason to believe that such a spectrum indicates the existence of a very extensive and very hot atmosphere surrounding the main body, or core, of the star in question. This particular star is remarkable in that it has undergone great changes in brilliancy and is located upon a background of nebulosity. The chances are strong that the star has rushed through the nebulosity with high rate of speed and that the resulting bombardment of the star has expanded and intensely heated its atmosphere.

There are the Wolf-Rayet stars, named from the French astronomers who discovered the first three of this class, whose spectra show a great variety of combinations of continuous spectrum and bright bands. We believe that the continuous spectrum in such a star comes from the more condensed central part, or core, and that the

There is an occasional star, like chi Carinae, whose spectrum consists almost wholly of bright lines, in general bearing no apparent relationship to the bright lines in the spectra of the gaseous nebulae except that the hydrogen lines are there, as they are almost everywhere. There is reason to believe that such a spectrum indicates the existence of a very extensive and very hot atmosphere surrounding the main body, or core, of the star in question. This particular star is remarkable in that it has undergone great changes in brilliancy and is located upon a background of nebulosity. The chances are strong that the star has rushed through the nebulosity with high rate of speed and

16 Pt

There is an occasional star, like chi Carinae, whose spectrum consists almost wholly of bright lines, in general bearing no apparent relationship to the bright lines in the spectra of the gaseous nebulae except that the hydrogen lines are there, as they are almost everywhere. There is reason to believe that such a spectrum indicates the existence of a very extensive and very hot atmosphere surrounding the main body, or core, of the star in question. This particular star is remarkable in that it has undergone great changes in brilliancy and is located upon a background of nebulosity. The chances are strong that the star has rushed through the nebulosity with high rate of speed and that the resulting bombardment of the star has expanded and intensely heated its atmosphere.

There are the Wolf-Rayet stars, named from the French astronomers who discovered the first three of this class, whose spectra show a great variety of combinations of continuous spectrum and bright bands. We believe that the continuous spectrum in such a star comes from the more condensed central

There is an occasional star, like chi Carinae, whose spectrum consists almost wholly of bright lines, in general bearing no apparent relationship to the bright lines in the spectra of the gaseous nebulae except that the hydrogen lines are there, as they are almost everywhere. There is reason to believe that such a spectrum indicates the existence of a very extensive and very hot atmosphere surrounding the main body, or core, of the star in question. This particular star is remarkable in that it has undergone great changes in brilliancy and is located upon a background of nebulosity. The chances are strong that the star has rushed through the nebulosity with high rate of speed and

16 Pt

There is an occasional star, like chi Carinae, whose spectrum consists almost wholly of bright lines, in general bearing no apparent relationship to the bright lines in the spectra of the gaseous nebulae except that the hydrogen lines are there, as they are almost everywhere. There is reason to believe that such a spectrum indicates the existence of a very extensive and very hot atmosphere surrounding the main body, or core, of the star in question. This particular star is remarkable in that it has undergone great changes in brilliancy and is located upon a background of nebulosity. The chances are strong that the star has rushed through the nebulosity with high rate of speed and that the resulting bombardment of the star has expanded and intensely heated its atmosphere.

There are the Wolf-Rayet stars, named from the French astronomers who discovered the first three of this class, whose spectra show a great variety of combinations of continuous spectrum and bright bands. We believe that the continuous spectrum in such a star comes from the more condensed central

#### 

**UPPERCASE** 

#### 

LOWERCASE

0123456789

0123456789 \$¢£¥€₫f

NUMBERS AND CURRENCY

Mathematical Symbols

Punctuation

Symbols